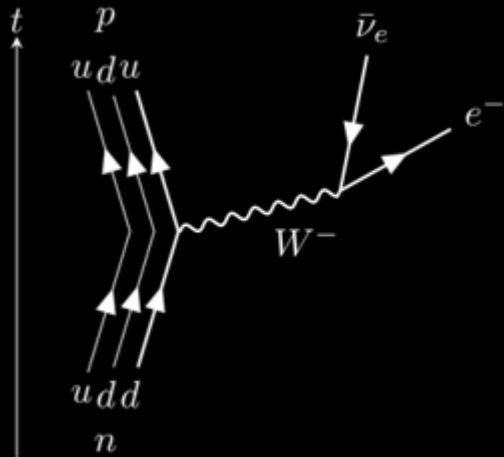




$0\nu\beta\beta$ decay searches with the CUORE and CUPID experiments

End of Year Seminars
September 17-18, Pavia
MANENTI Nicola

Massless Neutrinos: Standard Model



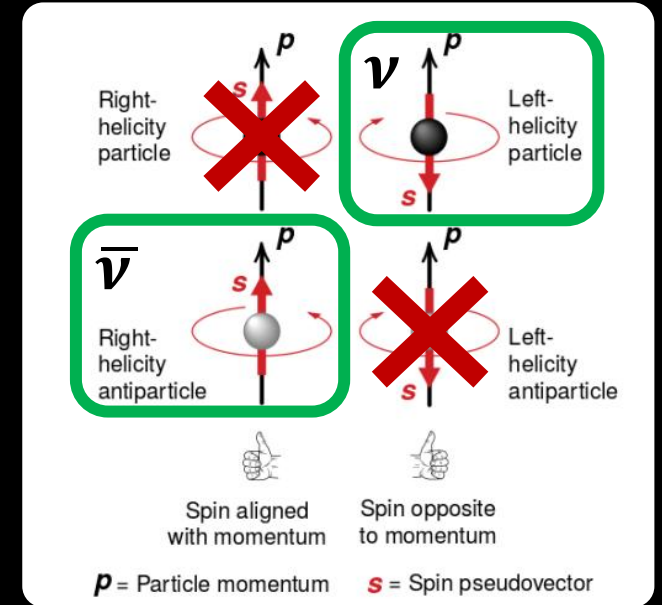
Neutrinos (ν) are fermions (spin = 1/2)
 → Dirac equation:

$$(i\gamma^\mu \partial_\mu - m)\nu = 0 \quad \begin{cases} i\gamma^\mu \partial_\mu \nu_L = m\nu_R \\ i\gamma^\mu \partial_\mu \nu_R = m\nu_L \end{cases}$$

No experimental proof of their mass
 → theory can be simplified

$$\begin{cases} i\gamma^\mu \partial_\mu \nu_L = 0 \\ i\gamma^\mu \partial_\mu \nu_R = 0 \end{cases}$$

Helicity measurements “prove” it



Nino the Neutrino

from “*Dal Sole alla Terra*”
 a CERN & INFN Production



Neutrino Oscillations



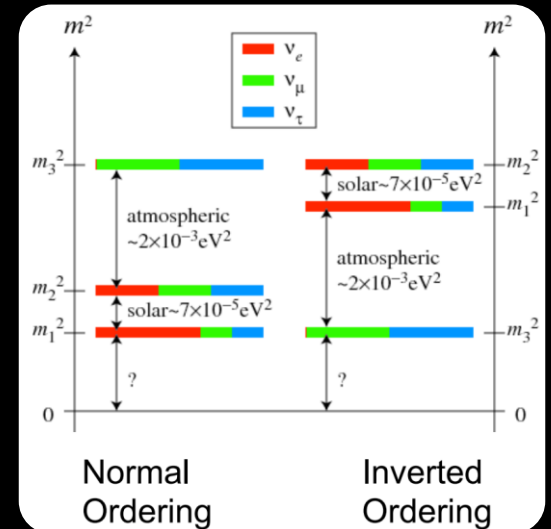
- ❖ neutrinos have different mass and flavor eigenbasis

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle$$

- ❖ at least two neutrinos need to be massive

$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2 \left(1.27 \frac{\Delta m^2 L}{E} \frac{[\text{eV}^2] [\text{km}]}{[\text{GeV}]} \right)$$

- ❖ oscillation experiments cannot probe the mass values, ambiguity in the mass ordering



Pontecorvo-Maki-Nakagawa-Sakata mixing matrix

$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Observable
Majorana phases
Dirac \mathcal{CP} phase

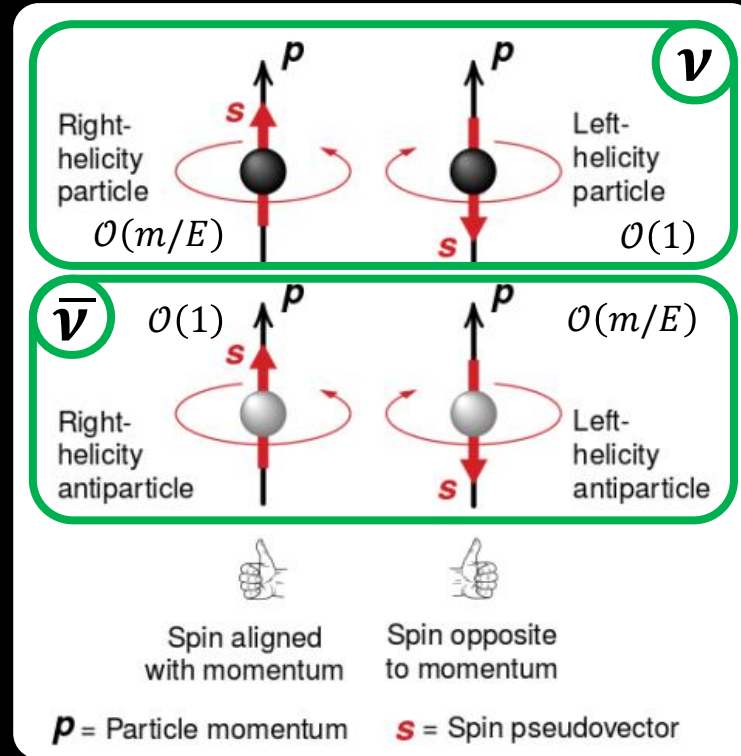
Massive Neutrinos and SM

1. Can we add a mass term to the neutrinos in the Standard Model?
2. What about the smallness of their mass?



P. Dirac

ν s are massive:
helicity \neq chirality



4 possible states:

$$\nu = \nu_L + \nu_R$$

$$\bar{\nu} = \bar{\nu}_L + \bar{\nu}_R$$

mass term:

$$m_D \bar{\nu}_L \nu_R + \text{h.c.}$$

$$\nu_{L/R} = P_{L/R} \nu = \frac{1 \mp \gamma_5}{2} \nu$$

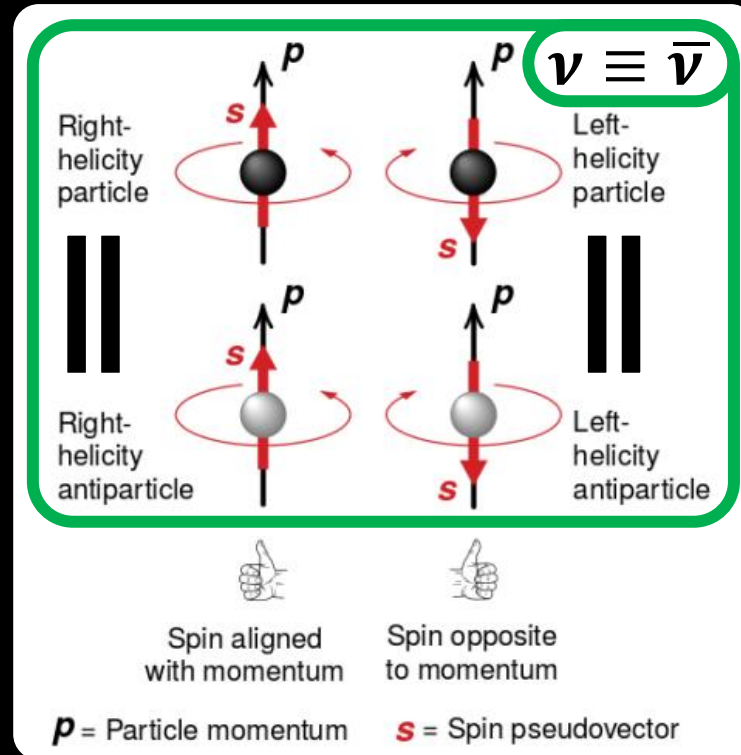
Massive Neutrinos and SM

1. Can we add a mass term to the neutrinos in the Standard Model?
2. What about the smallness of their mass?



E. Majorana

ν and $\bar{\nu}$ are different helicity states of the same particle



2 possible states:

$$\nu = \nu_L + \nu_L^C = \nu^C$$

mass terms:

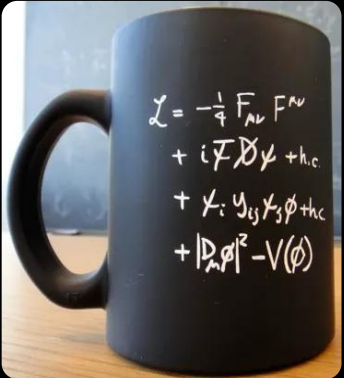
$$(m_L \nu_L^T C \nu_L + \text{h.c.})$$

$$(m_R \nu_R^T C \nu_R + \text{h.c.})$$

SM and Massive Neutrinos

Standard Model

$$m_\nu = 0$$



$$\mathcal{L} = + i\bar{\nu}_L \partial_\mu \gamma^\mu \nu_L + i\bar{\nu}_R \partial_\mu \gamma^\mu \nu_R$$

Dirac Mass

$$m_D \sim m_{\text{quark/lepton}}$$



$$- (m_D \bar{\nu}_R \nu_L + \text{h.c.})$$

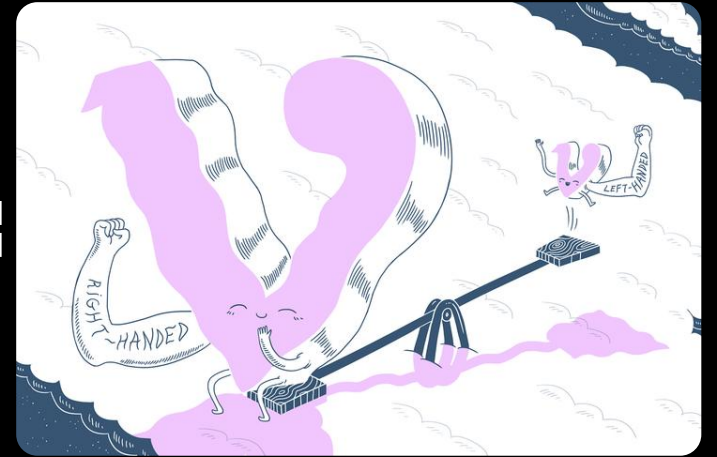
Majorana Mass

$$m_L = 0 \quad m_R \gg m_D$$



$$-\frac{1}{2} (m_L \bar{\nu}_L C \bar{\nu}_L^t + \text{h.c.}) - \frac{1}{2} (m_R \bar{\nu}_R C \bar{\nu}_R^t + \text{h.c.})$$

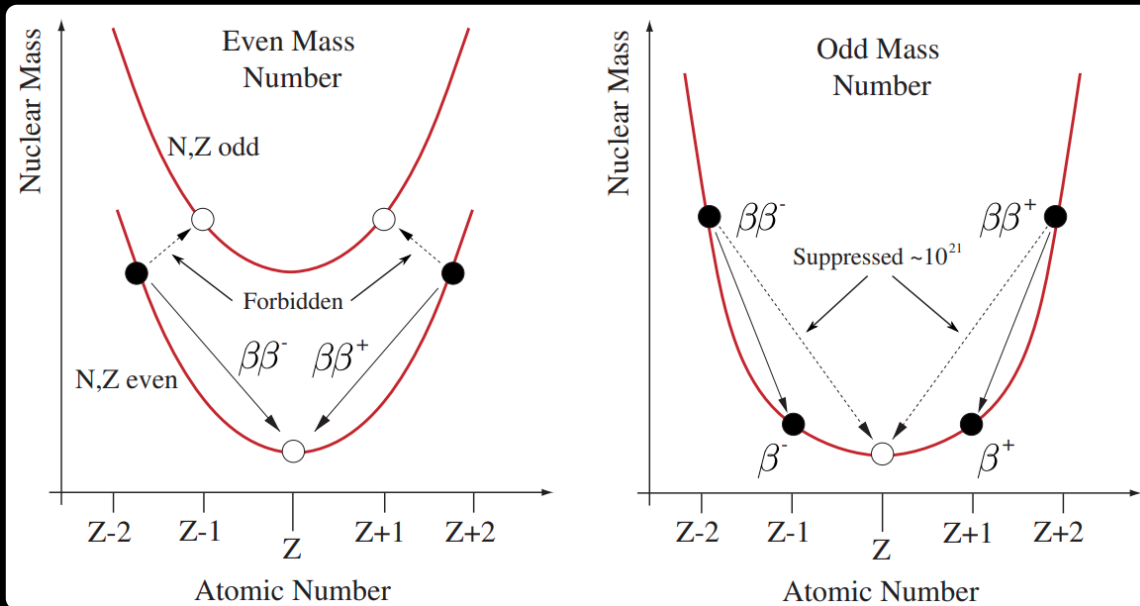
See-Saw Mechanism (e.g. one flavor)



$$m_1 \simeq \frac{m_D^2}{m_R} \ll m_D \quad m_2 \simeq m_R$$

$$\text{EW neutrino:} \quad \nu_L = i\nu_{1L} + \frac{m_D}{m_R} \nu_{2L}$$

$$\text{Sterile neutrino:} \quad (\nu_R)^c = -i \frac{m_D}{m_R} \nu_{1L} + \nu_{2L}$$

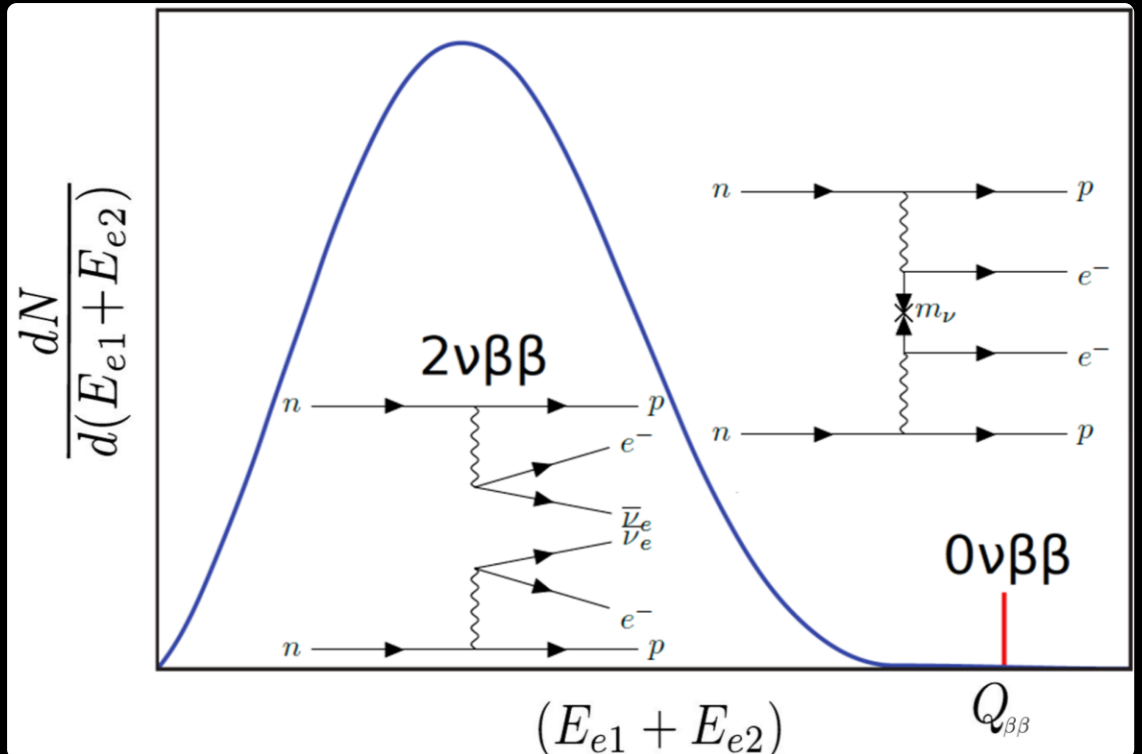


Double Beta Decay

- ❖ 2nd order weak transition allowed by the SM
- ❖ Introduced by M. Goeppert-Mayer in 1935
- ❖ Rare process: $T_{1/2}^{2\nu\beta\beta} \sim 10^{18-21}$ yr
- ❖ 35 isotope candidates

Neutrinoless DBD

- ❖ Beyond Standard Model counterpart of DBD
- ❖ Proposed by W.H. Furry in 1939
- ❖ Very rare process: $T_{1/2}^{0\nu\beta\beta} > 10^{24-26}$ yr



Sensitivity

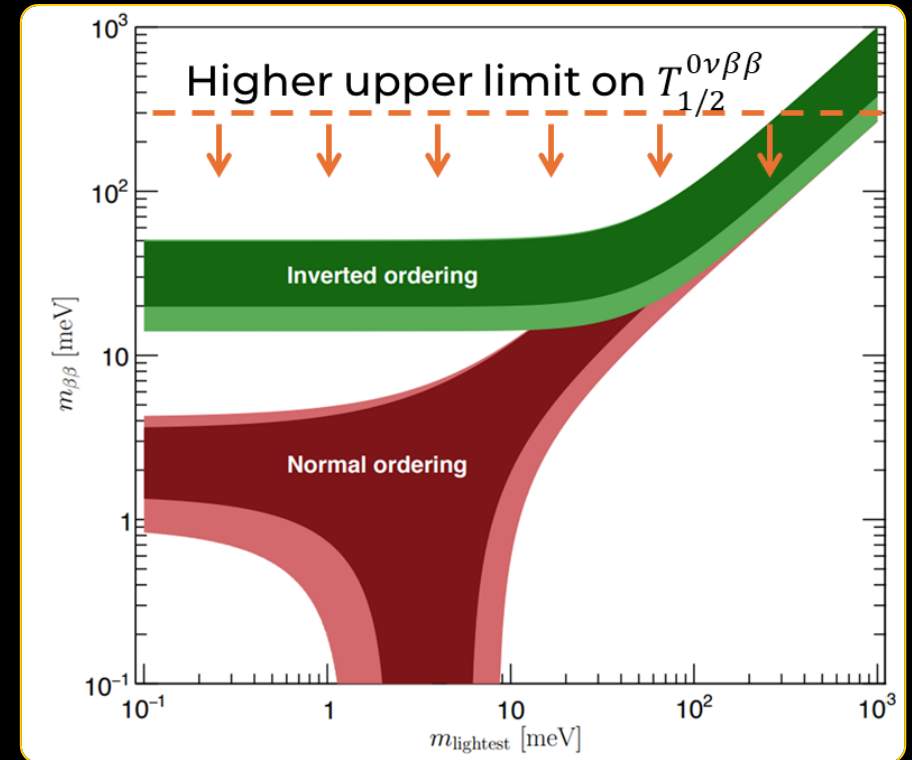
IF detected

- ❖ Matter-antimatter asymmetry
→ leptogenesis mechanism ($\Delta L = 2$)
- ❖ Neutrino mass scale and hierarchy
→ model dependent
- ❖ Majorana particles ($\nu \equiv \bar{\nu}$)

Observable

Effective
Majorana Mass

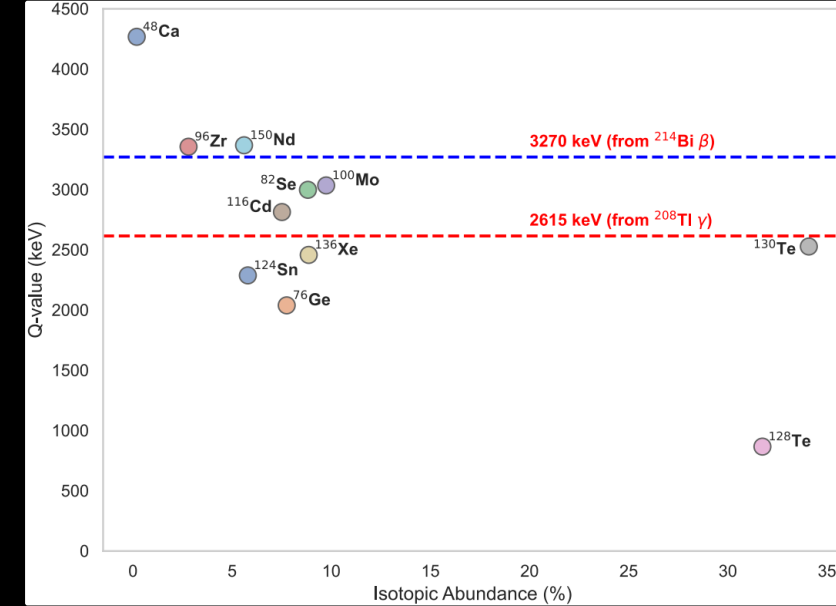
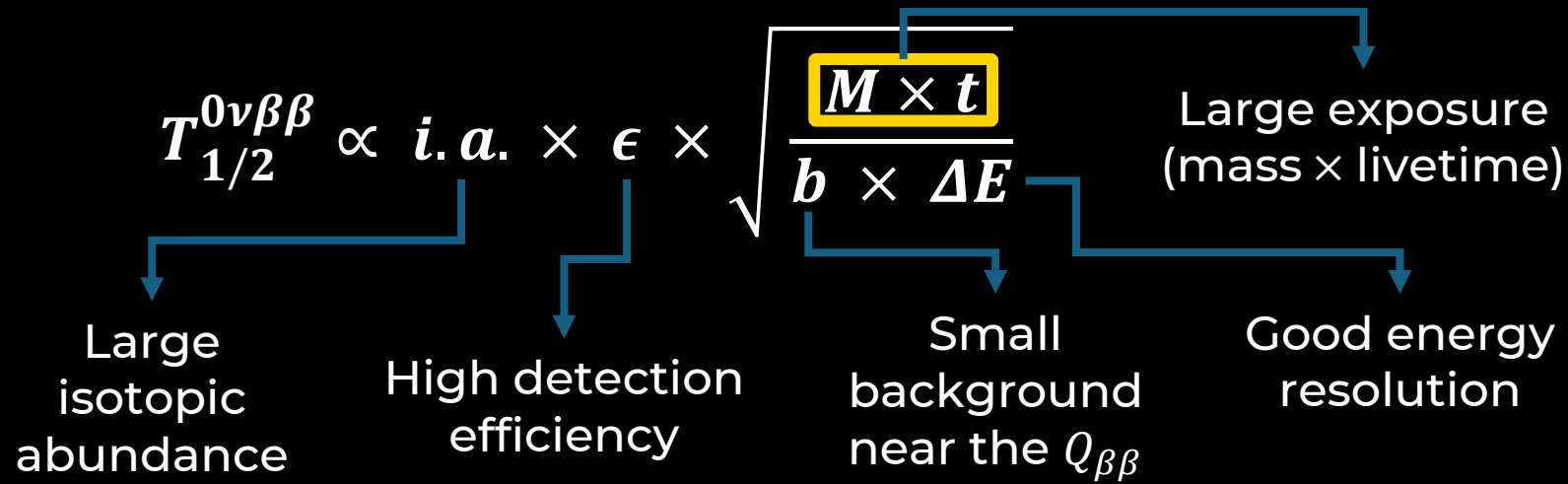
$$(m_{\beta\beta})^2 = \left| \sum_{i=1,2,3} U_{e,i}^2 m_i \right|^2$$



$$T_{1/2}^{0\nu\beta\beta} \propto \left[G^{0\nu} \times |M^{0\nu}|^2 \times m_{\beta\beta}^2 \right]^{-1}$$

\downarrow \downarrow \downarrow
 $0\nu\beta\beta$ half-life Phase Space Factor Nuclear Matrix Element

Detection Techniques



Granular Detectors

- ❖ Good Energy resolution
- ❖ Staging

→ Bolometers: CUORE, CUPID

→ Semiconductors: GERDA, LEGEND

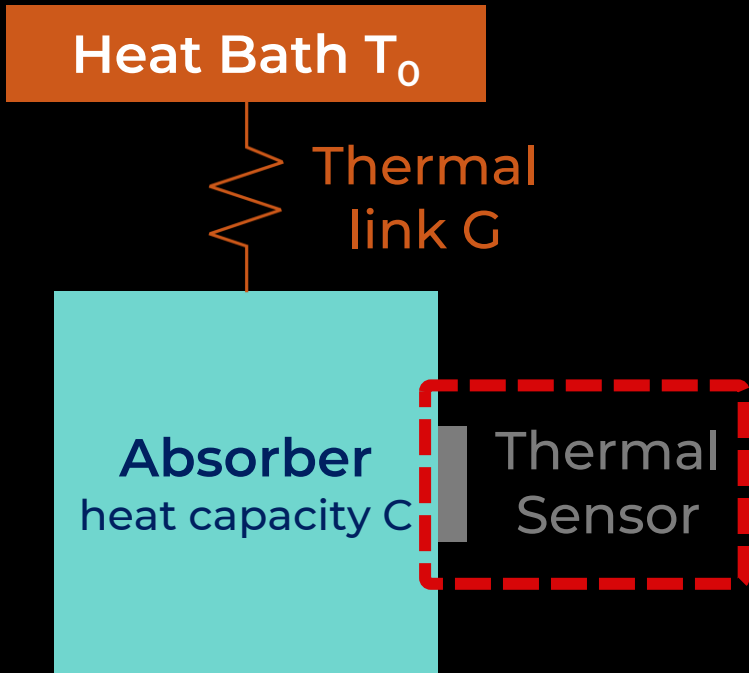
Monolithic Detector

- ❖ Shelf-shielding
- ❖ Scalability

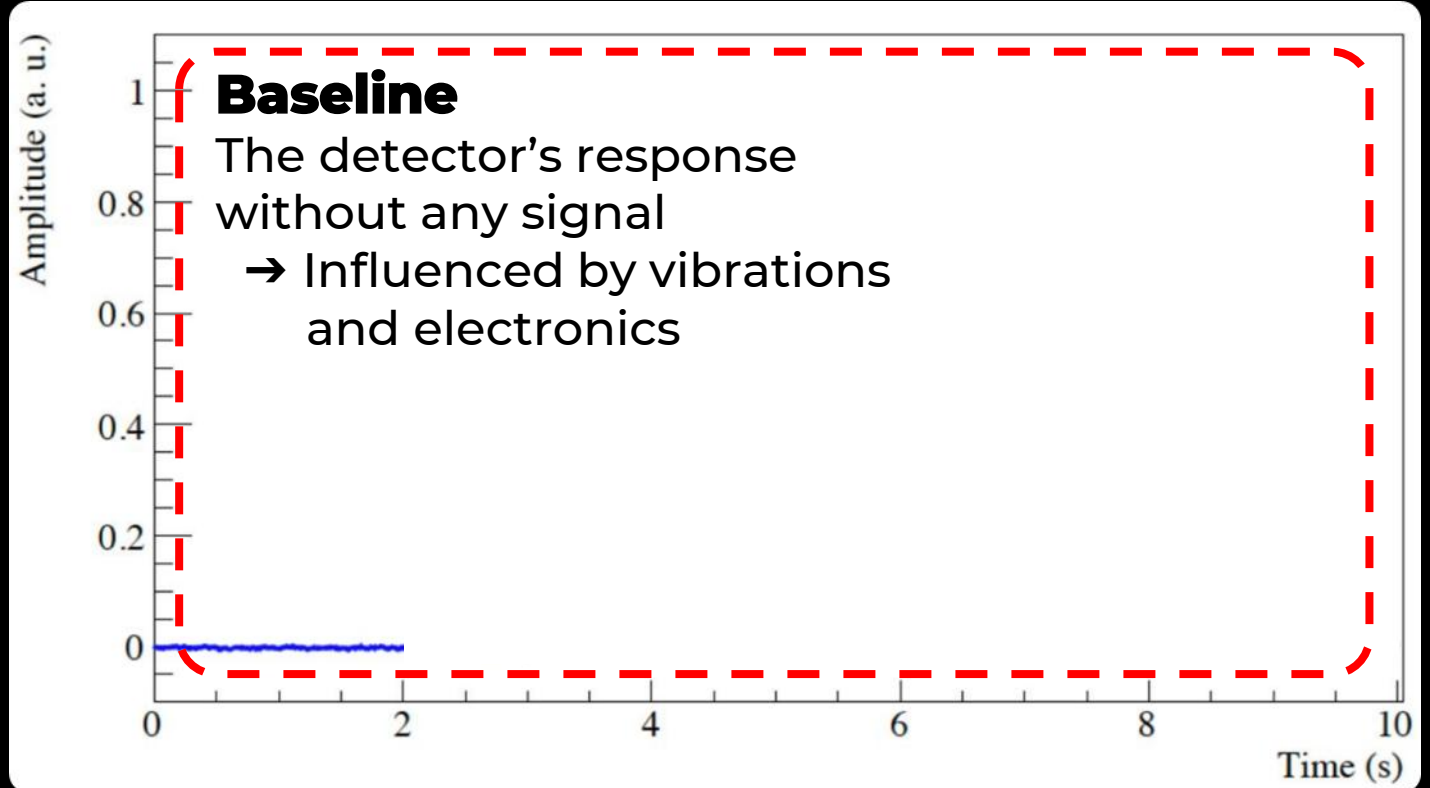
→ Scintillators: KamLAND-Zen, SNO

→ TPCs: nEXO, NEXT

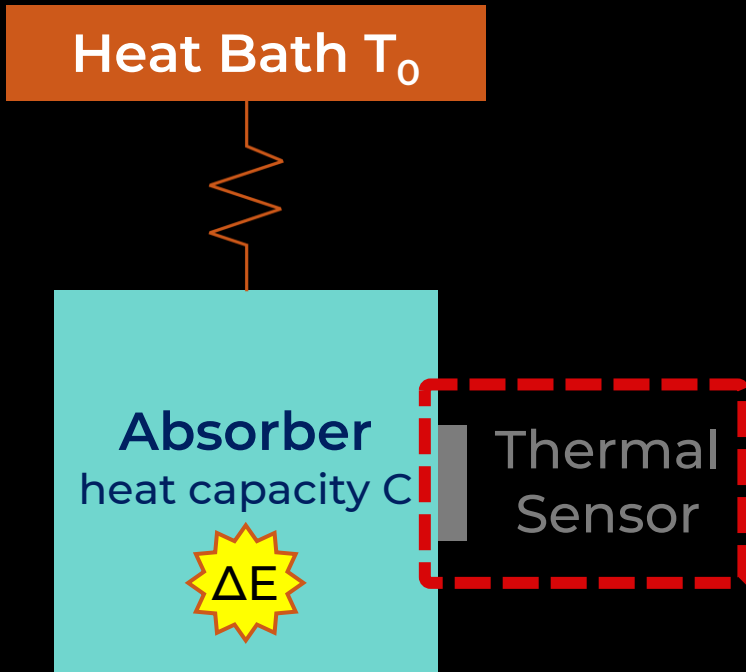
Cryogenic Calorimeters



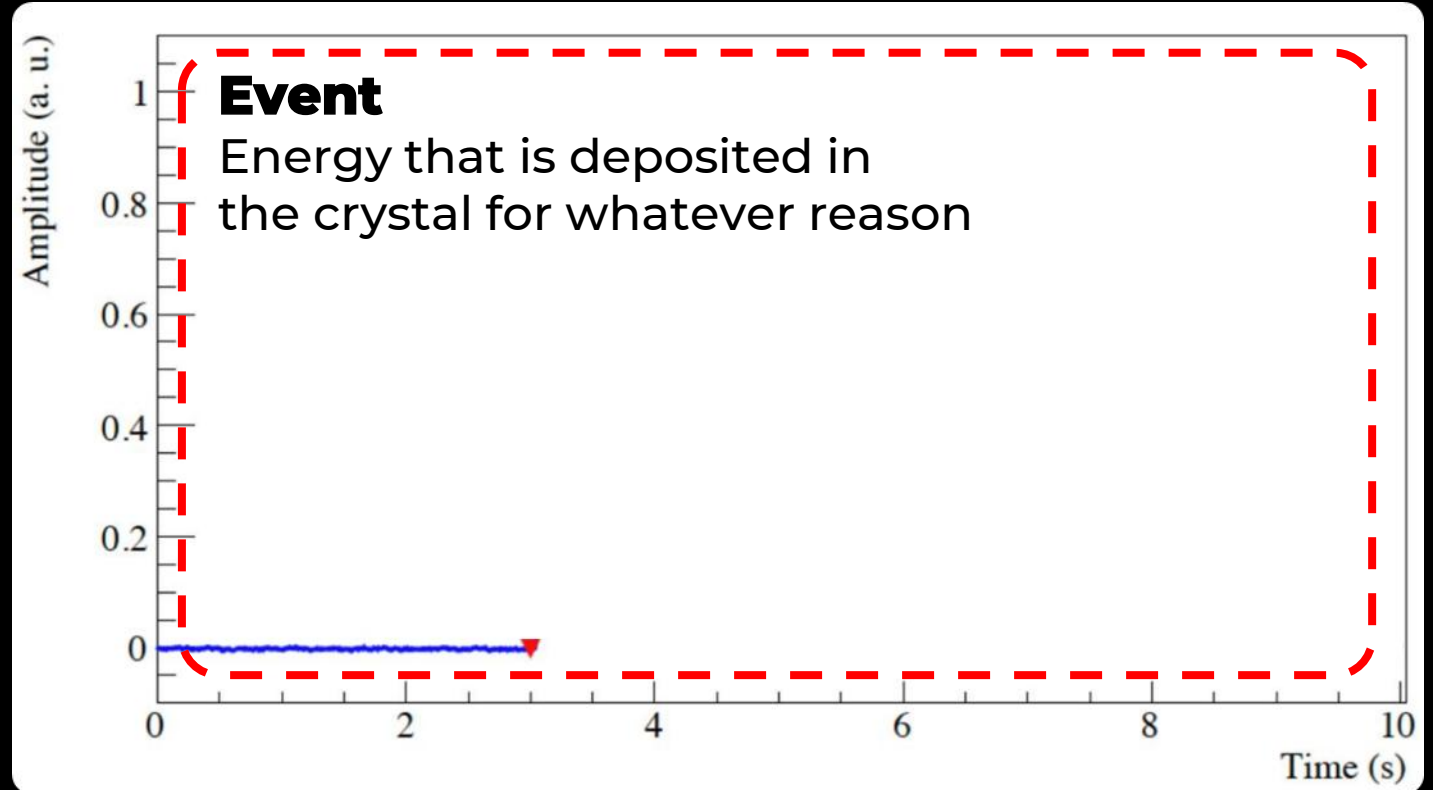
$$T = T_0$$



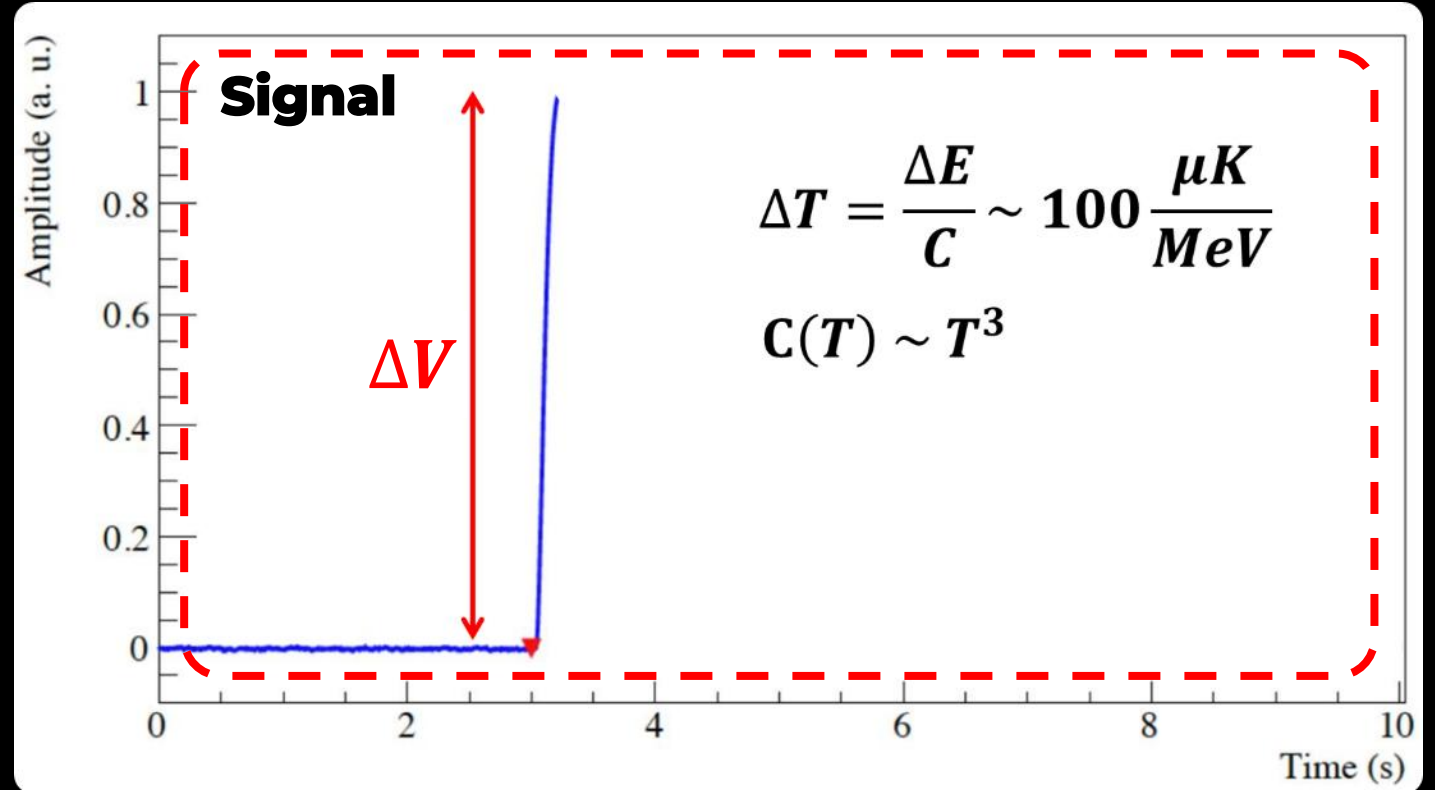
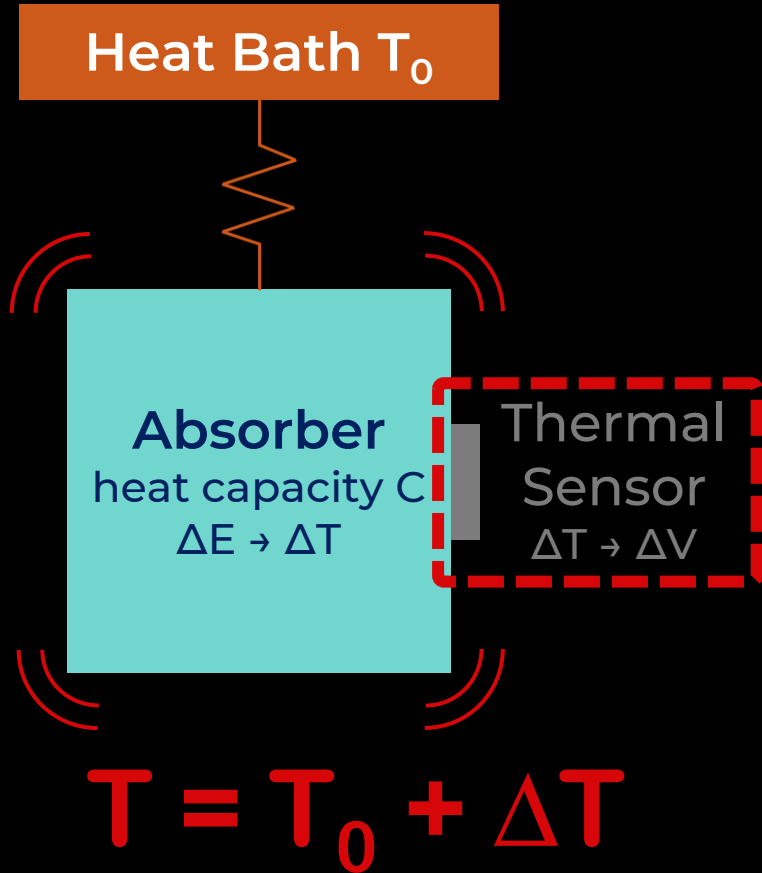
Cryogenic Calorimeters



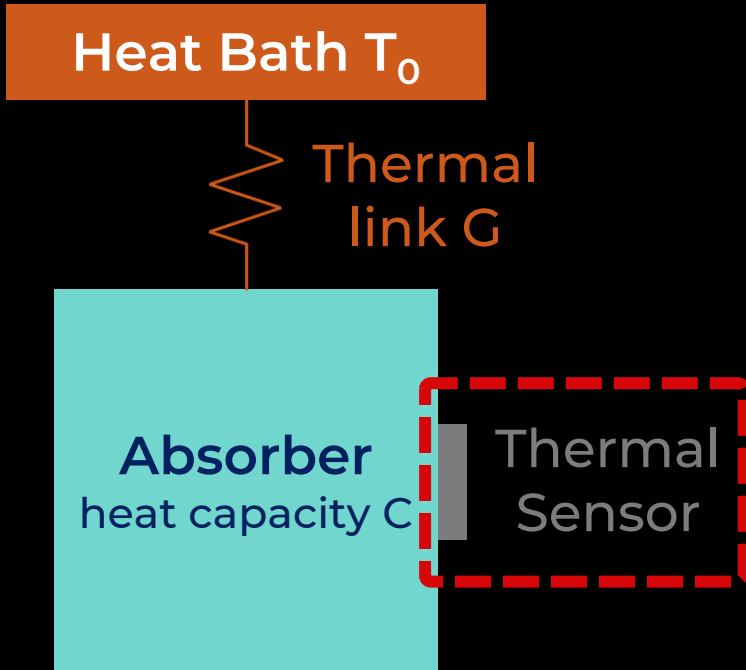
$$T = T_0$$



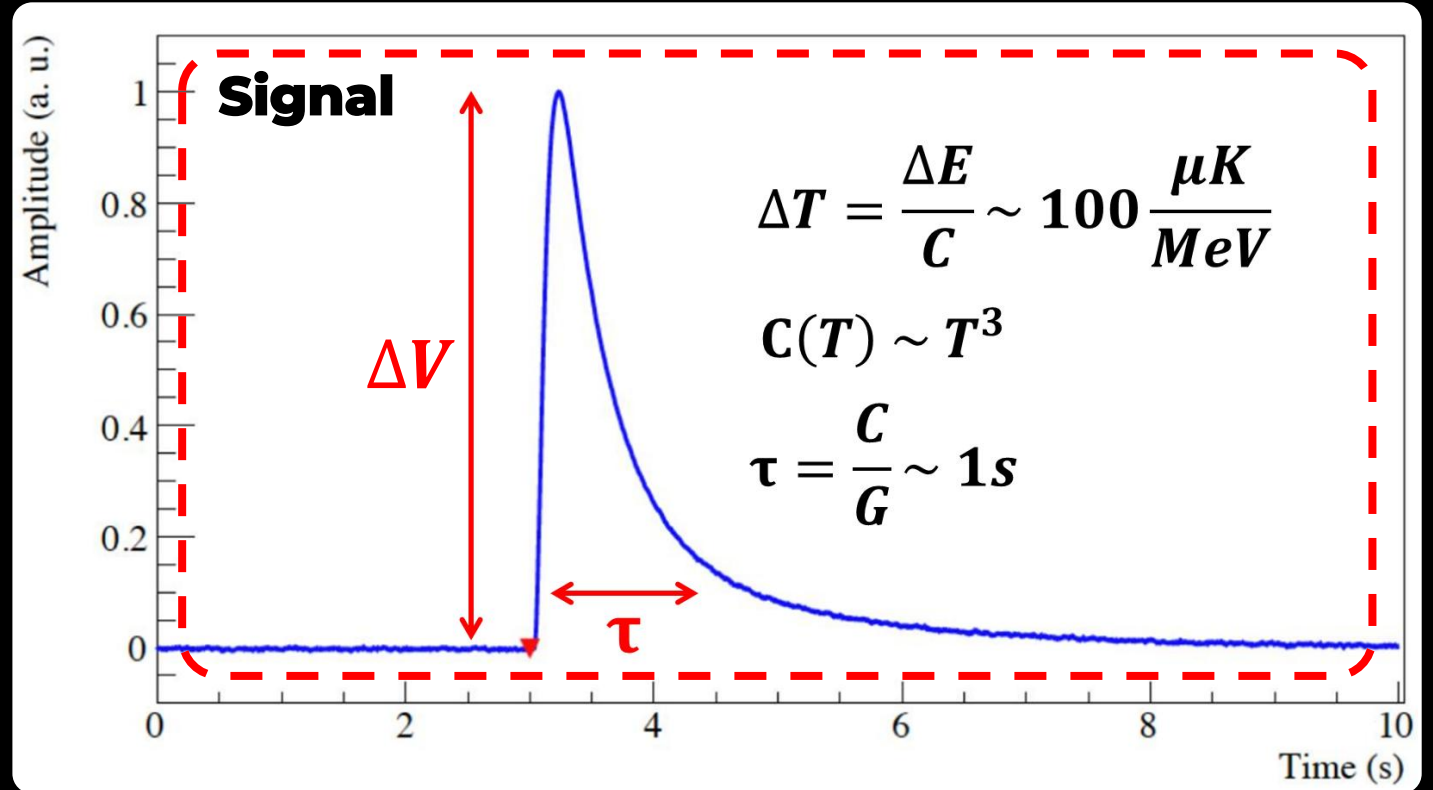
Cryogenic Calorimeters



Cryogenic Calorimeters



$$T = T_0$$





Challenges

Low Vibrations

- ❖ Noise diagnostic devices
- ❖ Decouple detectors and cryostat

Low Radioactivity

❖ Underground:

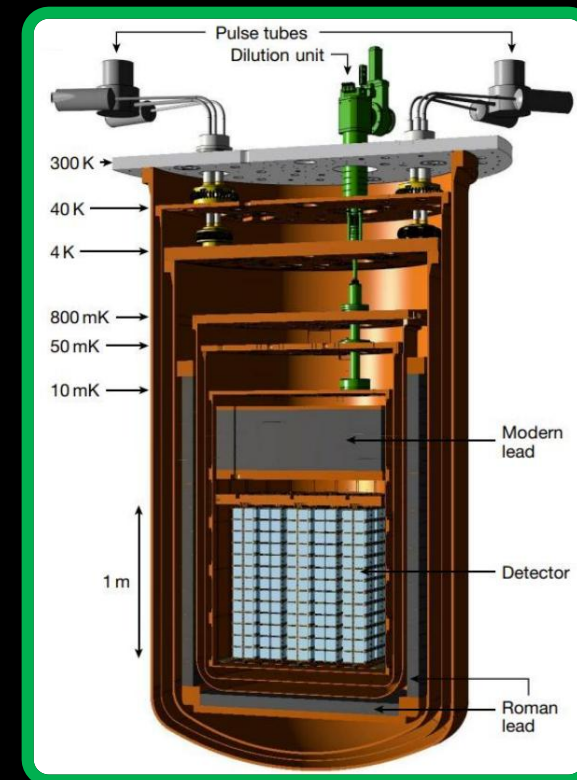
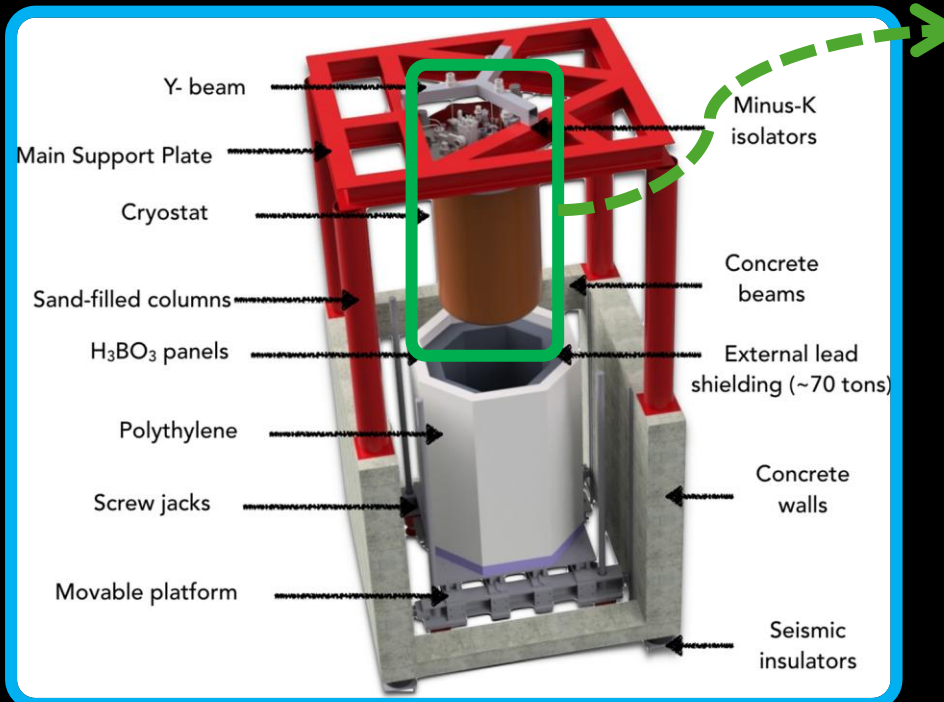
Average depth ~ 3600 m.w.e.

μ flux: $3 \times 10^{-8} \mu/s/cm^2$

n flux: $4 \times 10^{-6} n/s/cm^2 < 10 \text{ MeV}$

γ flux: $\sim 0.73 \gamma/s/cm^2 < 3 \text{ MeV}$

❖ External shields



Low Temperatures

❖ Mass < 4K: 15 tons

< 50mK: 3 tons

❖ Operating stably below 20mK for $\gg 1\text{yr}$

5x5x5 cm³ crystals

19 towers

13 floors

4 positions

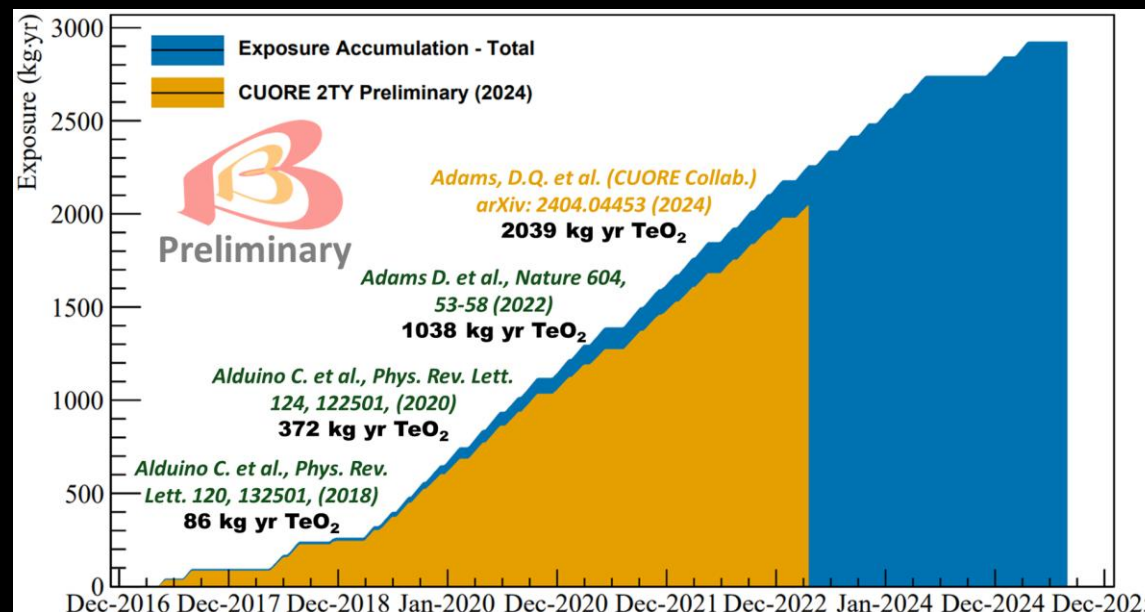


CUORE with

¹³⁰Te

34% (i.a.)
2517.5keV ($Q_{\beta\beta}$)

- ❖ 988 TeO₂ natural crystals → total ¹³⁰Te mass of 206kg
- ❖ Dynamic range: ~ keV → ~ MeV
- ❖ $\Delta E/E \sim 0.3\%$ @ $Q_{\beta\beta}$

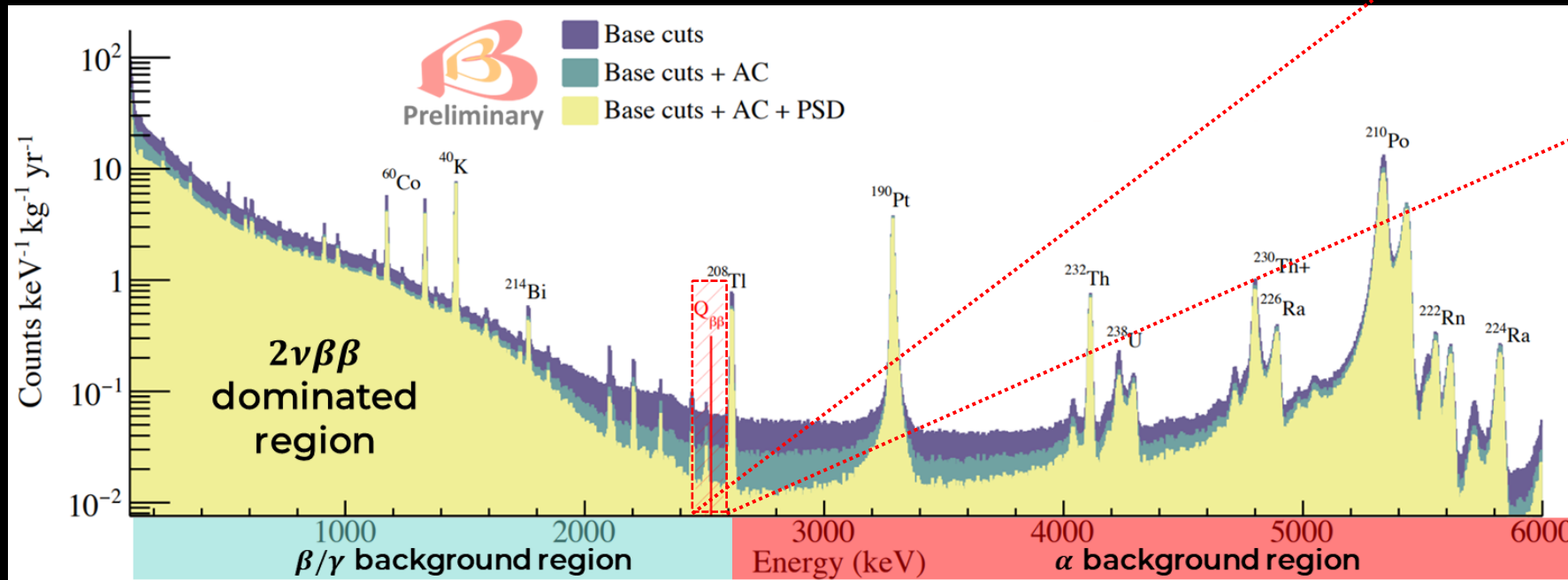
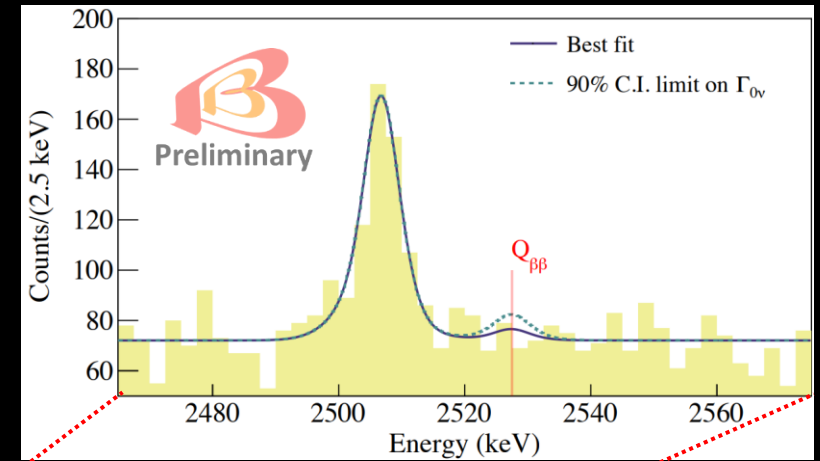


Cryogenic U
Observatory for Rare **E**vents

Goal: 3 tonnes year of exposure

CUORE - Results

- ❖ Data-driven background model
→ limits on other rare events
- ❖ Stable operation of a tonne-scale
milli-kelvin cryogenic calorimeter
- ❖ $T(^{100}\text{Te})_{1/2}^{0\nu\beta\beta} > 3.8 \times 10^{25} \text{ yr}$ (90% C.I.)



Base cuts
trigger, pile-up,
energy reconstruction

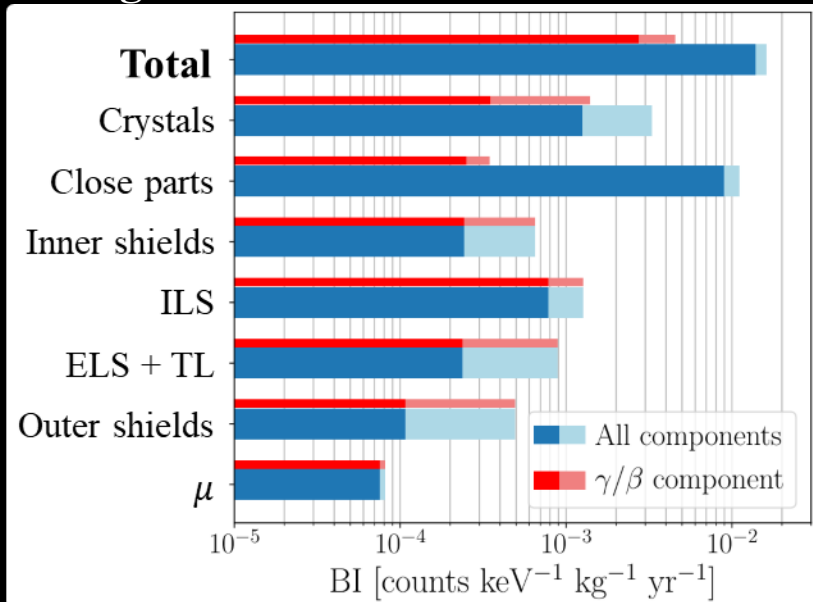
Anti-coincidence (AC)
only single-crystal events

Pulse shape discrimination (PSD)
only signal-like events

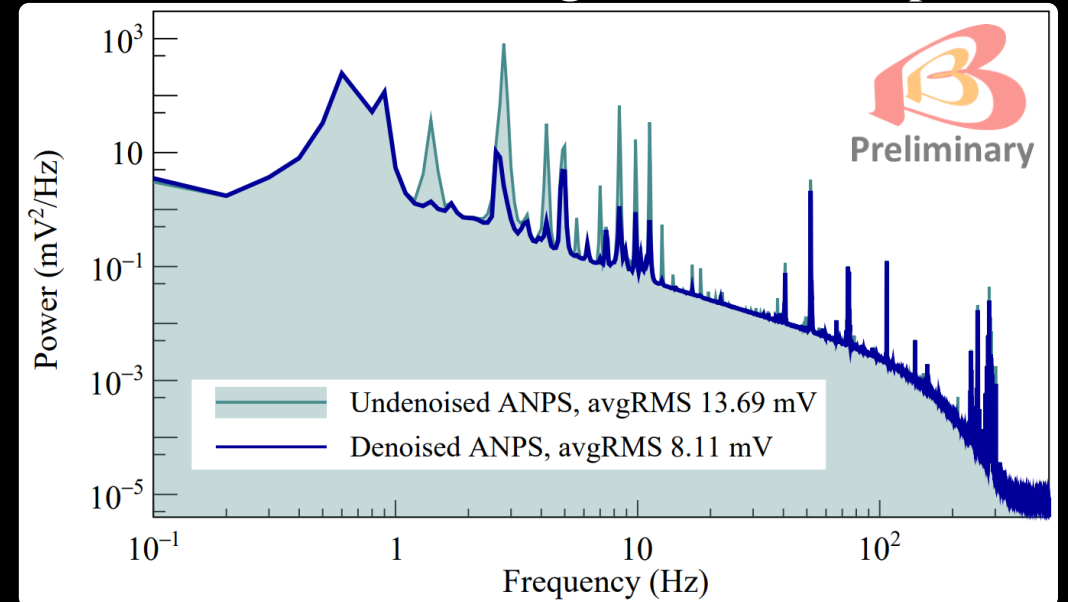
... and limitations

Vibrational noise and noise instabilities
in the \sim Hz range spoil the sensitivity
→ Denoising techniques

Background Index



Average Noise Power Spectra



CUORE is background-limited
by energy degraded α particles
→ Particle Identification

CUPID with

^{100}Mo

10% (i.a.)
3034keV ($Q_{\beta\beta}$)

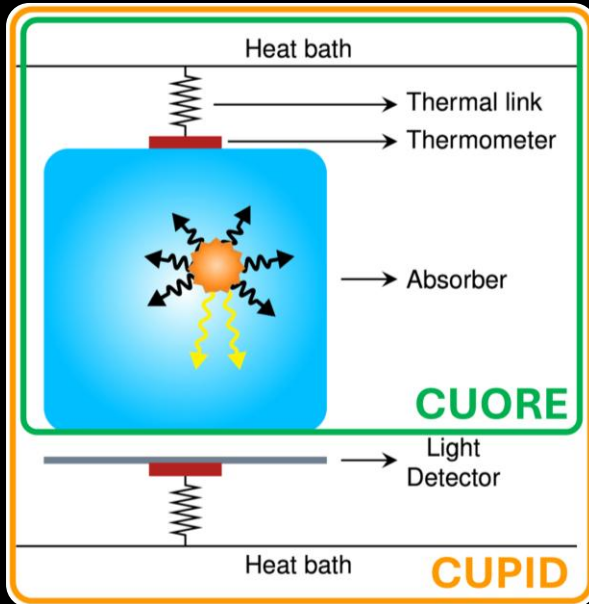
Favourable phase
space and matrix
element factors

4.5x4.5x4.5 cm³ crystals

57 towers

14 floors

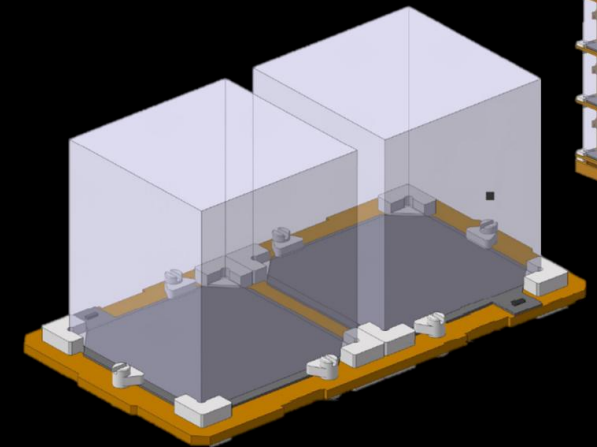
2 positions



Baseline design

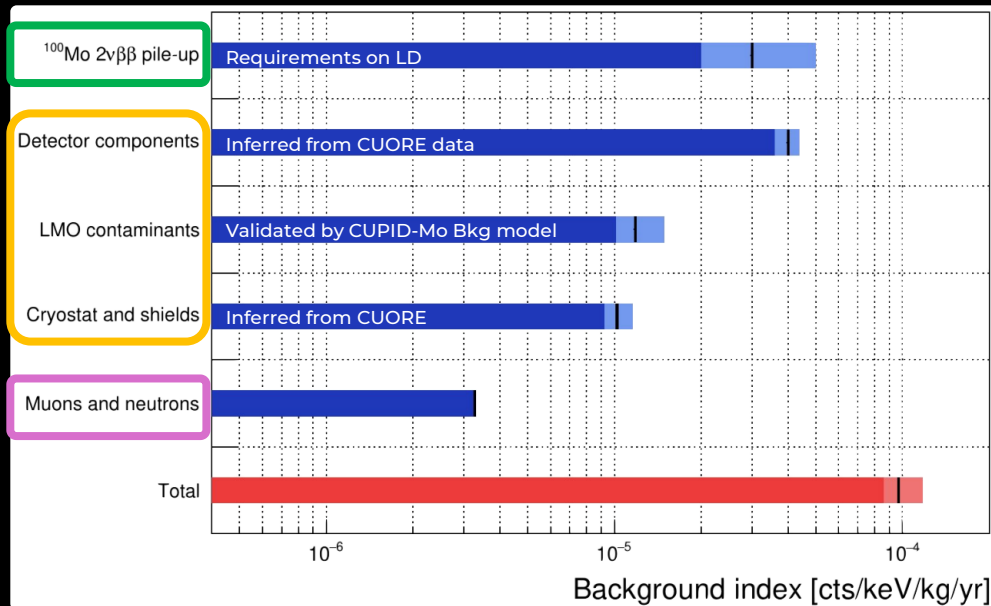
- ❖ Next-generation tonne scale experiment
- ❖ Will inherit CUORE's infrastructure
- ❖ Double readout strategy (heat and light)
- ❖ 1596 $\text{Li}_2^{100}\text{MoO}_4$ enriched crystals
→ 95% of ^{100}Mo , total mass of 240kg
- ❖ 1710 Ge-wafer Ligh Detectors
- ❖ $\Delta E < 5\text{keV}$ @ $Q_{\beta\beta}$

GOAL: $T_{1/2}^{0\nu\beta\beta} > 1.0 \times 10^{27} \text{ yr}$
(90% C.L. with 10 yr of livetime)



**CUORE Upgraded with
CUPID Particle IDentification**

CUPID – Background Budget



- ❖ $Q_{\beta\beta}$ far from the γ -background
- ❖ Lower noise and higher bandwidth electronics
 - pulse tubes upgrade and thermalisation optimisation
 - auxiliary environmental sensors for denoising
- ❖ Material selection, cleaning, shielding
- ❖ Delayed coincidence cuts (U/Th chains).



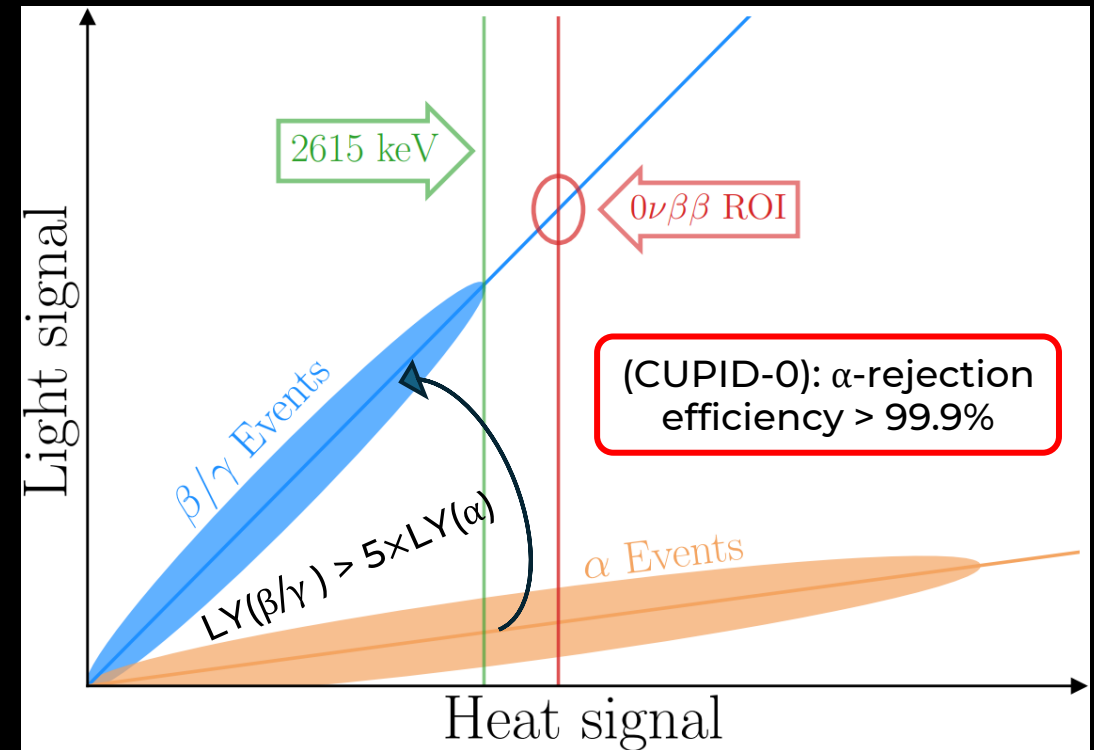
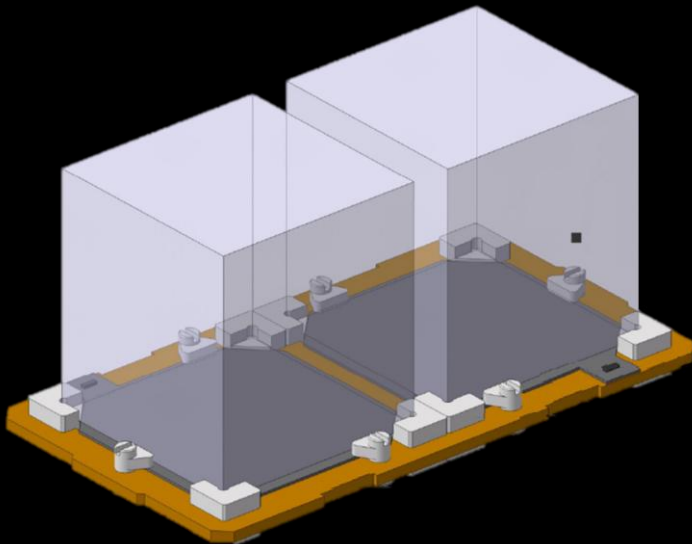
Muon tagger
veto detectors
and neutron
absorbers

$$T(^{100}\text{Mo})_{1/2}^{2\nu\beta\beta} = 7.1 \times 10^{18} \text{ yr}$$

Random coincidences with $2\nu\beta\beta$ events not so negligible
→ constraints on the LD timing resolution and SNR

LD - Requirements

- ❖ Particle Identification
 - $LY \sim 0.3 \text{ keV/MeV}$
 - Baseline resolution $< 100 \text{ eV RMS}$
- ❖ Particle Shape Discrimination
 - NTL amplification
 - Timing resolution $< 170 \mu\text{s}$



Neganov-Trofimov-Luke amplification

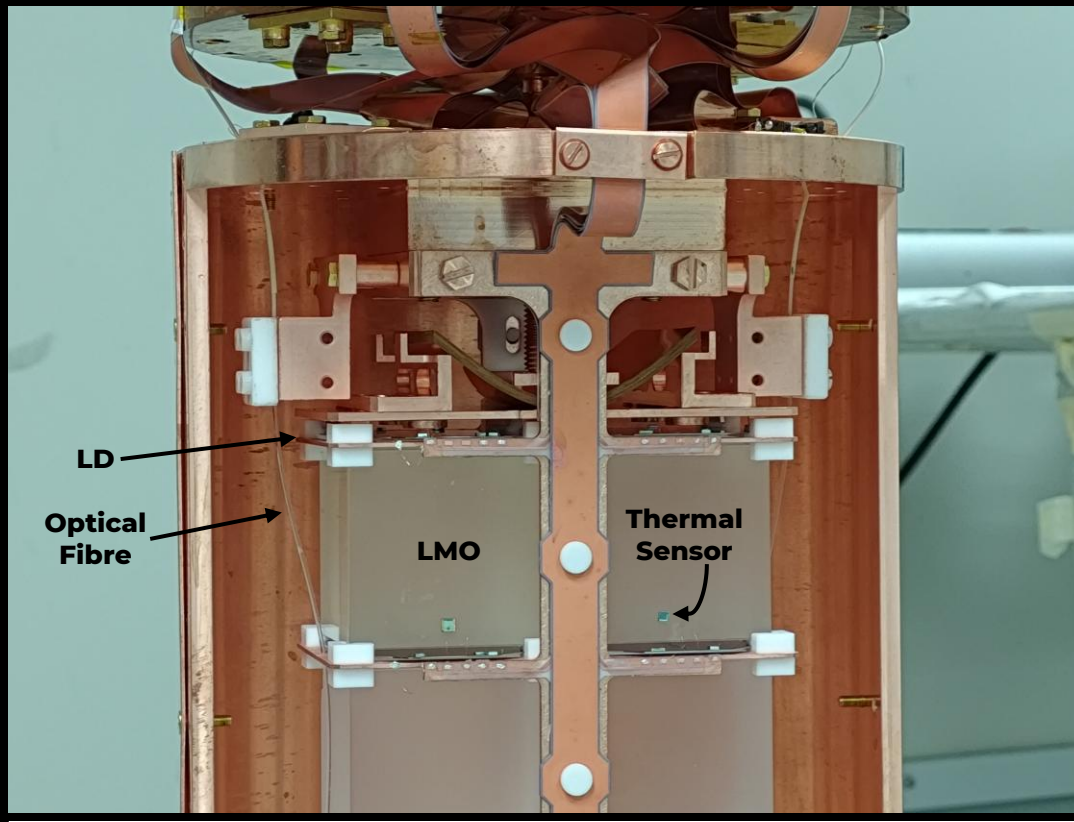
Electron-hole pairs created by light absorption are drifted by applying a voltage, thus producing additional heat

→ Charge accumulation is a problem

Optical Injection System

A system capable of injecting light pulses of a given wavelength to be absorbed by a group of LDs.

Baseline: (Light Source + Optical Fibre) x 57 Towers



Applications

- ❖ Pile-up ID efficiency monitoring
- ❖ LD stabilisation
- ❖ LD periodic regeneration
- ❖ Energy calibration

Requirements

- ❖ Multichannel
- ❖ Low radioactivity
- ❖ Low thermal load

Design

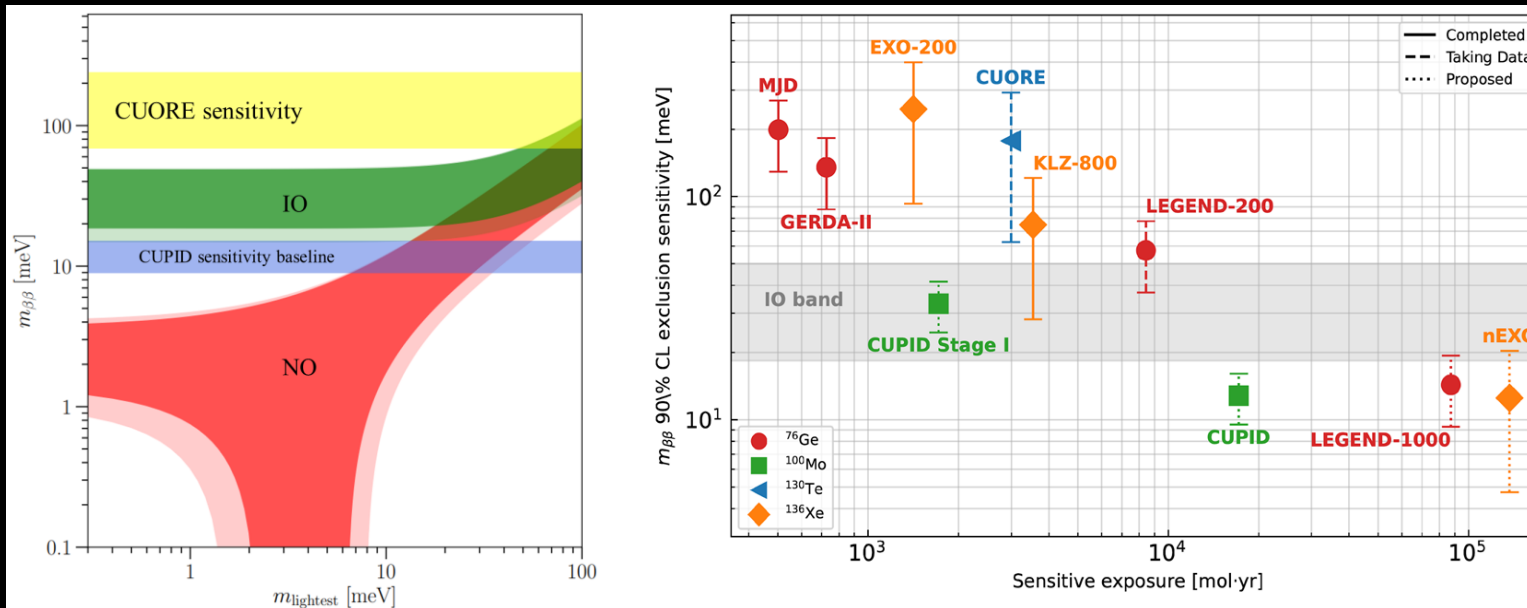
- ❖ Stable pulses
- ❖ Wavelength

Source

- ❖ Pulse width
- ❖ DAQ interface

Electronics

What's next !?



2026



2030

CUORE - 3 tonnes·yr limit
Cryogenic system upgrades
CUORE phase 2
→ higher sensitivity for
low energy signatures
→ test of the upgrades

CUPID Commissioning

THANK YOU FOR YOUR ATTENTION



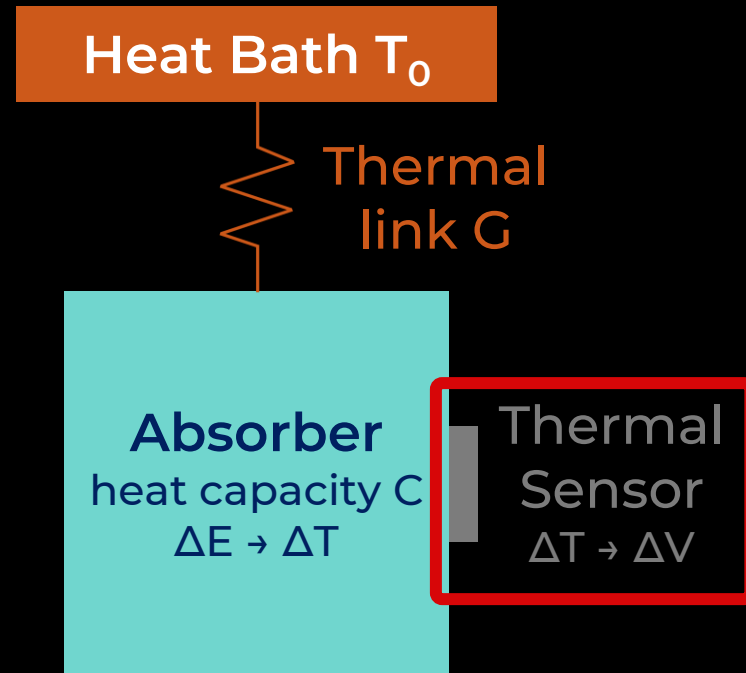
CUORE and CUPID collaborations, Collaboration Meeting, LNGS May 2025

Main References

- ❖ Toward the discovery of matter creation with neutrinoless $\beta\beta$ decay, Rev.Mod.Phys. 95 (2023) 2, 025002
- ❖ Search for Majorana neutrinos exploiting millikelvin cryogenics with CUORE, Nature 604 (2022) 7904, 53-58
- ❖ Optimization of the CUORE detector during the commissioning phase, Stefano Dell'Oro PhD Thesis
- ❖ With or without ν ? Hunting for the seed of the matter-antimatter asymmetry, arXiv: 2404.04453
- ❖ Sensitivity of the CUPID experiment to $0\nu\beta\beta$ decay of ^{100}Mo , arXiv:2504.14369
- ❖ CUPID, the Cuore upgrade with particle identification, Eur.Phys.J.C 85 (2025) 7, 737

Backup

Cryogenic Calorimeters - NTDs



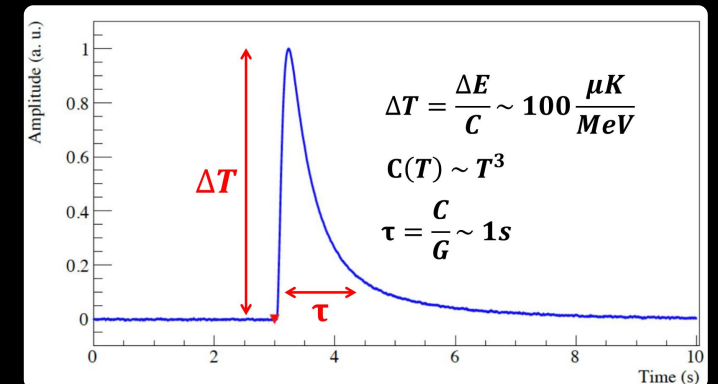
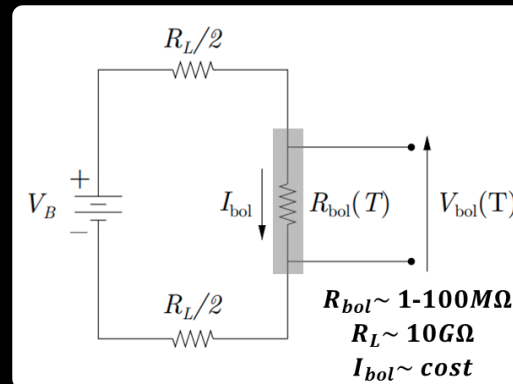
Neutron Transmutation Doped (Ge-NTD)

- Semiconductor Thermistor glued on the absorber
- Standard readout electronics
- Dynamic range from $\sim 10\text{keV}$ up to $\sim 10\text{MeV}$
- At $T \ll 10\text{K}$, resistivity is temperature dependent

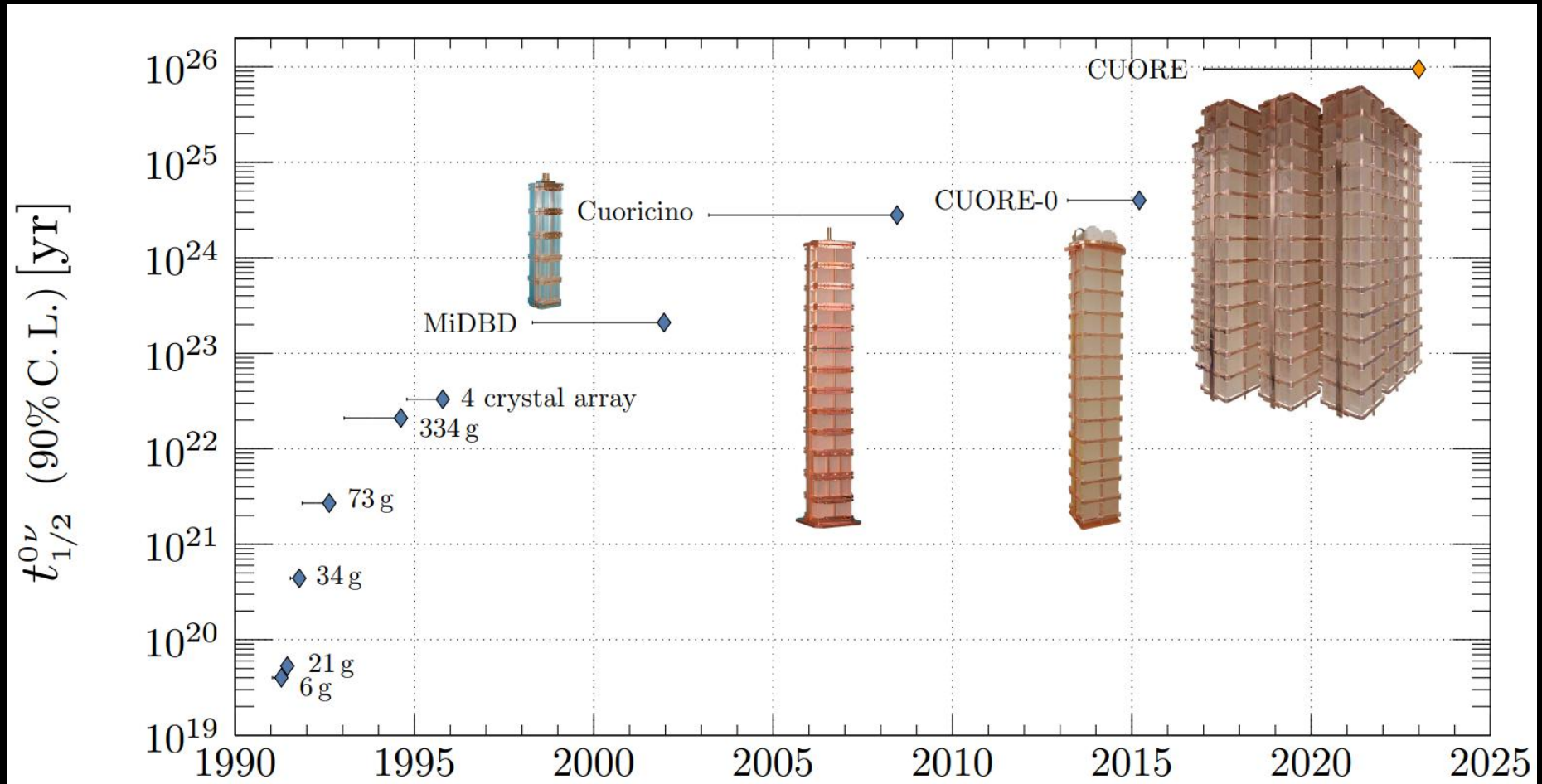
$$\rho = \rho_0 e^{\sqrt{T_0/T}}$$

where (ρ_0, T_0) depends on the doping level

- Sensible to mechanical vibrations and microphonism



Path toward CUORE



CUPID R&D Test

Dilution Cryostat

